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A Thermodynamically-Consistent Non-Ideal Stochastic Hard-Sphere Fluid

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A grid-free variant of the Direct Simulation Monte Carlo (DSMC) method is proposed, named the Isotropic DSMC (I-DSMC) method, that is suitable for simulating collisiondominated dense fluid flows. The I-DSMC algorithm eliminates all grid artifacts from the traditional DSMC algorithm and is Galilean invariant and microscopically isotropic. The stochastic collision rules in I-DSMC are modified to introduce a non-ideal structure factor that gives consistent compressibility, as first proposed in [Phys. Rev. Lett. 101:075902 (2008). The resulting Stochastic Hard Sphere Dynamics (SHSD) fluid is empirically shown to be thermodynamically identical to a deterministic Hamiltonian system of penetrable spheres interacting with a linear core pair potential, well-described by the hypernetted chain (HNC) approximation. We develop a kinetic theory for the SHSD fluid to obtain estimates for the transport coefficients that are in excellent agreement with particle simulations over a wide range of densities and collision rates. The fluctuating hydrodynamic behavior of the SHSD fluid is verified by comparing its dynamic structure factor against theory based on the Landau-Lifshitz Navier-Stokes equations. We also study the Brownian motion of a nano-particle suspended in an SHSD fluid and find a long-time power-law tail in its velocity autocorrelation function consistent with hydrodynamic theory and molecular dynamics calculations.

With the increased interest in nano- and micro-fluidics, it has become necessary to develop tools for hydrodynamic calculations at the atomistic scale [1, 2]. There are several issues present in microscopic flows that are difficult to account for in models relying on the continuum Navier-Stokes equations. Firstly, it is complicated to deal with boundaries and interfaces in a way that consistently accounts for the bidirectional coupling between the flow and (moving) complex surfaces or suspended particles. Furthermore, it is not trivial to include thermal fluctuations in Navier-Stokes solvers [3–5], and in fact, most of the time the fluctuations are not included even though they can be very important at instabilities [6] or in driving the dynamics of suspended objects [7, 8]. Finally, since the grid cell sizes needed to resolve complex microscopic flows are small, a large computational effort is needed even for continuum solvers. An alternative is to use particle-based methods, which are explicit and unconditionally stable and rather simple to implement. The

fluid particles are directly coupled to the microgeometry, for example, they directly interact with the beads of a polymer chain. Fluctuations occur naturally and the algorithm is designed to give the correct spatio-temporal correlations.

Several particle methods have been described in the literature. The most accurate but also most expensive is molecular dynamics (MD) [9], and several coarse-grained models have been developed, such as dissipative particle dynamics (DPD) [10] and multi-particle collision dynamics (MPCD) [11, 12], each of which has its own advantages and disadvantages [13]. Our method, first proposed in Ref. [14], is based on the Direct Simulation Monte Carlo (DSMC) algorithm of Bird [15]. The key idea behind DSMC is to replace deterministic interactions between the particles with stochastic momentum exchange (collisions) between nearby particles. While DSMC is usually viewed as a kinetic Monte Carlo algorithm for solving the Boltzmann equation for a low-density gas, it can also be viewed as an alternative to the expensive MD in cases where an approximate (coarse-grained) treatment of the molecular transport is appropriate. The stochastic treatment of collisional treatment makes the algorithm much simpler and faster than MD, while preserving the essential ingredients of fluctuating hydrodynamics: local momentum conservation, linear momentum exchange on length scales comparable to the particle size, and a similar fluctuation spectrum.

Being composed of point particles, the DSMC fluid has no internal structure, has an ideal gas equation of state (EOS), and is thus very compressible. As a consequence, the density fluctuations in DSMC are significantly larger than those in realistic liquids. Furthermore, the speed of sound is small (comparable to the average speed of the particles) and thus subsonic (Mach number less than one) flows are limited to relatively small Reynolds numbers¹. Efforts have been undertaken to develop coarse-grained models that have greater computational efficiency than brute-force MD and that have a non-ideal EOS, such as the Lattice-Boltzmann (LB) method [16], DPD [17], MPCD [18, 19]. The Consistent Boltzmann Algorithm (CBA) [20, 21], as well as algorithms based on the Enskog equation [22, 23], have demonstrated that DSMC fluids can have dense-fluid compressibility, however, they did not achieve thermodynamic consistency between the equation of state and the fluid structure.

In this paper we describe a generalization of the traditional DSMC algorithm suitable for dense

¹ For a low-density gas the Reynolds number is Re = M/K, where $M = v_{flow}/c$ is the Mach number, and the Knudsen number $K = \lambda/L$ is the ratio between the mean free path λ and the typical obstacle length L. This shows that subsonic flows can only achieve high Re flows for small Knudsen numbers, i.e., large numbers of DSMC particles.

fluid flows. By a dense fluid we mean a fluid where the mean free path is small compared to the typical inter-atomic distance. As a first step, we introduce a grid-free *Isotropic DSMC* (I-DSMC) method that eliminates all grid artifacts from traditional DSMC, notably the lack of Galilean invariance and non-isotropy. The I-DSMC fluid is still an *ideal* fluid just like the traditional DSMC fluid, that is, it has an the equation of state of an ideal gas and does not have an internal structure as do liquids. Secondly, by biasing the collision kernel in I-DSMC to only allow stochastic collisions between approaching particles, we obtain the *Stochastic Hard-Sphere Dynamics* (SHSD) algorithm that is thermodynamically consistent (i.e., the direct calculation of compressibility from density fluctuations agrees with the density derivative of pressure). The SHSD algorithm is related to previous algorithms for solving the Enskog kinetic equation [22, 23], and can be viewed as a more-efficient variable-diameter stochastic modification of the traditional hard-sphere molecular dynamics [24].

In the SHSD algorithm randomly chosen pairs of approaching particles that lie less than a given diameter of each other undergo collisions as if they were hard spheres of diameter equal to their actual separation. The SHSD fluid is shown to be non-ideal, with structure and equation of state equivalent to that of a fluid mixture where spheres effectively interact with a repulsive linear core pairwise potential. We theoretically demonstrate this correspondence at low densities. Remarkably, we numerically find that this effective interaction potential, similar to the quadratic core potential used in many DPD variants, is valid at all densities. Therefore, the SHSD fluid, as DPD, is *intrinsically* thermodynamically-consistent, while non-ideal MPCD is only *numerically* thermodynamically-consistent for tuned choices of the parameters [18, 19].

The equivalence of the structure of the SHSD fluid with the linear core penetrable-sphere fluid enables us to use the Hypernetted Chain (HNC) approximation, as recommended in Ref. [25], to obtain theoretical estimates for the pair correlation and static structure factors that are in excellent agreement with the numerical results. These further enable us to use the Enskog-like kinetic theory developed in Ref. [26] to obtain excellent theoretical estimates of the transport properties of the SHSD fluid that are also shown to be in excellent agreement with numerics even at relatively high densities. This makes the SHSD fluid a rare example in which such simple but accurate theoretical estimates can be obtained.

We numerically demonstrate that the hydrodynamics of the SHSD fluid is consistent with the equations of fluctuating hydrodynamics when the appropriate equation of state taken into account. Specifically, we compare the dynamic structure factor with that obtained from the linearized fluctuating Navier-Stokes equations. We also calculate the velocity autocorrelation function (VACF)

for a large hard-sphere bead suspended in an SHSD fluid, demonstrating the existence of long-time tails as predicted by hydrodynamics and found in molecular dynamics simulations. The tail is found to be in quantitative agreement with theory at lower densities, but a discrepancy is found at higher densities, possibly due to the strong structuring of the SHSD fluid.

We begin by introducing a grid-free variant of the DSMC algorithm in Section I. This Isotropic DSMC algorithm simulates a stochastic particle system where particles closer than a particle diameter collide with a certain rate. By biasing the collision kernels to favor head-on collisions of particles, as in the hard-sphere fluid, we obtain a non-ideal stochastic fluid in Section II. We develop an Enskog-like kinetic theory for this Stochastic Hard Sphere Dynamics (SHSD) system in Section II A, which requires as input the pair correlation function. In Section II B we discover that the SHSD fluid is thermodynamically consistent with a fluid of penetrable linear core spheres, and use this equivalence to compute the pair correlation function of the SHSD fluid using the HNC approximation. In Section III we show several results, including a comparison between theory and numerics for the transport coefficients and for the dynamic structure factor, as well as a study of the hydrodynamic tails in the velocity autocorrelation of a bead suspended in an SHSD fluid.

I. ISOTROPIC DSMC

The traditional DSMC algorithm [15] starts with a time step where particles are propagated advectively, $\mathbf{r}'_i = \mathbf{r}_i + \mathbf{v}_i \Delta t$, and sorted into a grid of cells. Then, a certain number $N_{coll} \sim \Gamma_{sc}N_c(N_c-1)\Delta t$ of stochastic collisions are executed between pairs of particles randomly chosen from the N_c particles inside the cell, where the collision rate Γ_{sc} is chosen based on kinetic theory. The conservative stochastic collisions exchange momentum and energy between two particles i and j that is not correlated with the actual positions of the particles. Typically the probability of collision is made proportional to the magnitude of the relative velocity $v_r = |\mathbf{v}_{ij}|$ by using a conventional rejection procedure.

Traditional DSMC suffers from several grid artifacts, which become pronounced when the mean free path becomes comparable to the DSMC cell size. Firstly, the method is not Galilean invariant unless the grid of cells is shifted randomly before each collision step, as typically done in the MPCD/SRD algorithm [11, 12] for the same reason. This shifting is trivial in a purely particle simulation with periodic boundary conditions, but it causes implementation difficulties when boundaries are present and also in particle-continuum hybrids. Furthermore, traditional DSMC, unlike MD, is not microscopically isotropic and does not conserve angular momentum, leading to

an anisotropic stress tensor. Instead of trying to work around these grid artifacts, as done for MPCD in Refs. [18, 19], we have chosen to modify the traditional DSMC algorithm to make it grid-free.

To ensure isotropy, all particle pairs within a collision diameter D (i.e., overlapping particles if we consider the particles to be spheres of diameter D) are considered as potential collision partners even if they are in neighboring cells. In this way, the grid is only used as a tool to find neighboring particles efficiently, but does not otherwise affect the properties of the resulting stochastic fluid. Such a grid-free DSMC variant, which we will call the Isotropic Direct Simulation Monte Carlo (I-DSMC) method, is suitable for hydrodynamics of dense gases (where the mean free path is comparable or even smaller than D), unlike the original DSMC which targets the dilute limit. It is important to point out, however, that the I-DSMC is not meant to be a replacement for traditional DSMC for rare gas flows. In particular, the computational efficiency is reduced by a factor of 2-4 over traditional DSMC due to the need to search neighboring cells for collision partners in addition to the current cell. This additional cost is not justified at low densities, where the grid artifacts of traditional DSMC are small. Furthermore, the I-DSMC method is not intended as a solver for the Boltzmann equation, which was the primary purpose of traditional DSMC [27]. Rather, in the limit of small time steps, the I-DSMC method simulates the following stochastic particle system: Particles move ballistically in-between collisions. While two particles i and j are less than a diameter D apart, $r_{ij} \leq D$, there is a probability rate $\chi D^{-1}K_c(\boldsymbol{v}_{ij}, \boldsymbol{r}_{ij})$ for them to collide and change velocities without changing their positions, where K_c is some function of the relative position and velocity of the pair, and the dimensionless cross-section factor χ sets the collisional frequency. After the collision, the pair center-of-mass velocity does not change, ensuring momentum conservation, while the relative velocity is drawn from a probability density $P_c(\boldsymbol{v}_{ij}^{'}; \boldsymbol{v}_{ij}, \boldsymbol{r}_{ij})$, such that energy is conserved, $\|\boldsymbol{v}_{ij}^{'}\| = \|\boldsymbol{v}_{ij}\|$.

Once the pre- and post-collision kernels K_c and P_c are specified, the properties of the resulting I-DSMC fluid are determined by the cross-section factor χ and the density (hard-sphere volume fraction) $\phi = \pi N D^3/(6V)$, where N is the total number of particles in the simulation volume V. Compare this to the deterministic hard-sphere fluid, whose properties are determined by the volume fraction ϕ alone. It is convenient to normalize the collision kernel K_c so that for an ideal gas with a Maxwell-Boltzmann velocity distribution the average collisional rate would be χ times larger than that of a gas of hard spheres of diameter D at low densities, $\phi \ll 1$. Two particular choices for the collision kernels which we use in practice are:

Traditional DSMC collisions (Traditional I-DSMC ideal fluid), for which the probability of collision is made proportional to the magnitude of the relative velocity $v_{rel} = ||v_{ij}||$, $K_c = 3v_{rel}/4$. The post-collisional relative velocity is randomized so as to mimic the average distribution of collision impact parameters in a low-density hard-sphere gas, specifically, in three dimensions the relative velocity is rotated uniformly independent of r_{ij} [15]. We use this fluid mainly for comparison with traditional DSMC.

Maxwell collisions (Maxwell I-DSMC ideal fluid), for which $K_c = 3\sqrt{k_B T_0/m}/(2\sqrt{\pi})$ independent of the relative velocity, where T_0 is an equilibrium temperature. For simplicity, the post-collisional relative velocity is randomly rotated as in classic DSMC. This kernel is not realistic and may lead to unphysical results in cases where there are large density and temperature gradients, however, it is computationally most efficient since there is no rejection based on relative velocity. We therefore prefer this kernel for problems where the temperature dependence of the transport properties is not important, and what we will typically mean when we say I-DSMC without further qualification.

With a finite time step, the I-DSMC method can be viewed as a time-driven kinetic Monte Carlo algorithm for solving the Master Equation that describes the stochastic particle system. Unlike the Boltzmann equation, this Master Equation has a mollified collision kernel with a finite compact support D [26, 28], unlike the singular kernel in the Boltzmann equation. The traditional DSMC method also mollifies the collision kernel by considering particles within the same cell of size L_c as collision partners. This cell L_c however has nothing to do with the actual molecular diameter σ , in fact, for low densities it is typically much larger than σ , but still smaller than the mean free path. The molecular properties enter in traditional DSMC only in the form of collisional cross-sections $\sim \sigma^2$. In light of this, for rare gas flows, the collision diameter D in I-DSMC should be considered the equivalent of the cell length L_c , and not σ . Traditional DSMC is designed to reproduce a collision rate per particle per unit time equal to the Boltzmann rate, $\Gamma_B(\sigma) = C\sigma^2$, where C is a constant independent of σ . The I-DSMC method is designed to reproduce a collision rate

$$\Gamma_{I-DSMC} = \chi \Gamma_B(D) = \chi CD^2,$$

and therefore by choosing

$$\chi = \chi_B = \left(\frac{\sigma}{D}\right)^2$$

we get $\Gamma_{I-DSMC} = \Gamma_B(\sigma)$. Therefore, if I-DSMC is used as a replacement for traditional DSMC in order to simulate the transport in a low-density gas of hard-sphere of diameter σ , the collision

diameter D should be chosen to be some fraction of the mean free path (say, $D \approx \lambda/4 \gg \sigma$), and the cross-section prefactor set to $\chi_B \sim (\sigma/\lambda)^2 \ll 1$. At higher densities χ_B starts becoming comparable to unity and thus it is no longer possible to separate the kinetic and collisional time scales as assumed in traditional DSMC.

A. Performing Stochastic Collisions

In I-DSMC, stochastic collisions are processed at the end of every particle time step of duration Δt , and in-between collisions each particle i is streamed advectively with constant velocity v_i . At the end of each time step, we need to randomly and without bias choose pairs of overlapping particles for collision, given the current configuration of the system. This can be done, as in traditional DSMC, using a rejection Monte Carlo technique. Specifically, we need to choose a large number $N_{tc}^{(tot)} = \Gamma_{tc}^{(tot)} N_{pairs} \Delta t$ of trial collision pairs, and then accept the fraction of them that are actually overlapping as collision candidates. Here N_{pairs} is the number of possibly-overlapping pairs, for example, as a first guess one can include all pairs, $N_{pairs} = N(N-1)/2$. If the probability of overlap for one of the trial pairs is p_{geom} , then $p_{geom}N_{pairs}$ candidate collisions will be accepted, and thus the probability for choosing one of the overlapping pairs as a collision candidate is simply $\Gamma_{tc}^{(tot)} \Delta t$. If the probability of accepting a candidate pair ij for an actual collision is $p_{ij}^{(acc)}$ and Δt is sufficiently small, then the probability rate to actually collide i and j while they are overlapping approaches $\Gamma_{ij} = p_{ij}^{(acc)} \Gamma_{tc}^{(tot)}$. The goal is to choose the trial collision frequency $\Gamma_{tc}^{(tot)}$ and $p_{ij}^{(acc)}$ such that $\Gamma_{ij} = \chi D^{-1}K_c(v_{ij}, r_{ij})$.

The efficiency of the algorithm is increased if the probability of accepting trial collisions is increased. In order to increase p_{geom} , one should reduce N_{pairs} to be closer to the number of actually overlapping pairs (making N_{pairs} linear instead of quadratic in N), ideally, one would build a list of all the overlapping pairs. This is however expensive, and a reasonable compromise is to use collision cells similarly to what is done in classical DSMC and also MD algorithms. Namely, the spatial domain of the simulation, typically a unit periodic cell, is divided into cells of length $L_c \gtrsim D$, and for each cell a linked list \mathcal{L}_c of all the particles in that cell is maintained. All pairs of particles that reside in the same or neighboring cells are considered as potential collision partners, and here we include the cell itself in its list of neighboring cells, i.e., each cell has 3^d neighbors, where d is the spatial dimension. To avoid any spatial correlations (inhomogeneity and non-isotropy), trial collision pairs should be chosen at random one by one. This would require first choosing a pair of neighboring cells with the correct probability, and then choosing a particle

from each cell (rejecting self-collisions). This is rather expensive to do, especially at lower χ , when few actual collisions occur at each time step, and we have therefore chosen to use a method that introduces a small bias each time step, but is unbiased over many time steps. Specifically, we visit the cells one by one and for each cell c we perform $N_{tc}^{(c)} = \Gamma_{tc}^{(c)} N_c N_p \Delta t$ trial collisions between one of the N_c particles in that cell and one of the N_p particles in the 3^d neighboring cells. Here $\Gamma_{tc}^{(c)}$ is a local trial collision rate and it may depend on the particular cell c under consideration. Note that each of the $N_c(N_c - 1)$ trial pairs ij where both i and j are in cell c may be chosen twice (self collisions are rejected). Similarly, each of the $N_c(N_p - N_c)$ trial pairs where i and j are in different cells c and c' may be chosen twice, once when each of the cells c and c' is considered. Also note that it is important not to visit the cells in a fixed order during every time step. Unlike in traditional cells, where cells are independent of each other and can be visited in an arbitrary order (even in parallel), in I-DSMC it is necessary to ensure isotropy by visiting the cells in a random order, different every time step.

For the Maxwell pre-collision kernel, once a pair of overlapping particles i and j is found a collision is performed without additional rejection, therefore, we set $\Gamma_{tc}^{(c)} = \chi D^{-1} K_c/2 = 3\chi D^{-1} \sqrt{k_B T_0/m}/(4\sqrt{\pi}) = \text{const}$, and we have divided by two because of the double counting of each trial pair. For the traditional pre-collision kernel, and, as we shall see shortly, the SHSD pre-collision kernel, additional rejection based on the relative velocity v_{ij} is necessary. As in the traditional DSMC algorithm, we estimate an upper bound for the maximal value of the pre-collision kernel $K_c^{(max)}$ among the pairs under consideration and set $\Gamma_{tc}^{(c)} = \chi D^{-1} K_c^{(max)}/2$. We then perform an actual collision for the trial pair ij with probability

$$p_{ij}^c = K_c(\boldsymbol{v}_{ij}, \boldsymbol{r}_{ij}) / K_c^{(max)},$$

giving the correct collision probability for every overlapping pair of particles. For the traditional pre-collision kernel $K_c^{(max)} = 3v_{rel}^{(max)}/4$, where $v_{rel}^{(max)}$ is as tight an estimate of the maximum relative speed as possible. In the traditional DSMC algorithm $v_{rel}^{(max)}$ is a global bound obtained by simply keeping track of the maximum particle speed v_{max} and taking $v_{rel}^{(max)} = 2v_{max}$ [15]. In I-DSMC, we obtain a local estimate of $v_{rel}^{(max)}$ for each cell c that is visited, thus increasing the acceptance rate and improving efficiency.

Algorithm 1 specified the procedure for performing collisions in the I-DSMC method. The algorithm is to a large degree collision-kernel independent, and in particular, the same algorithm is used for ideal and non-ideal stochastic fluids. As already explained, the size of the cells should be chosen to be as close as possible but still larger than the particle diameter D. The time step should

be chosen such that a typical particle travels a distance $l_{\Delta t} \approx \Delta t \sqrt{kT_0/m} \sim D\delta t$, where δt is a dimensionless time step, which should be kept reasonably smaller than one. It is also important to ensure that each particle does not on average undergo more than one collision per time step; we usually keep the number of collisions per particle per time step less than one half.

Algorithm 1: Processing of stochastic collisions between overlapping particles at a time-step in the I-DSMC method.

- 1. Sample a random permutation of the cell numbering \mathcal{P} .
- 2. Visit the cells one by one in the random order given by \mathcal{P} . For each cell c, do the following steps if the number of particles in that cell $N_c > 0$, otherwise move on to the next cell.
- 3. Build a list \mathcal{L}_1 of the N_c particles in the cell and at the same time find the largest particle speed in that cell v_1^{max} . Also keep track of the second largest speed in that cell v_2^{max} , which is an estimate of the largest possible speed of a collision partner for the particle with speed v_1^{max} .
- 4. Build a list of the N_p particles in the set of 3^d cells that neighbor c, including the cell c itself and respecting the proper boundary conditions. Also update v_2^{max} if any of the potential collision partners not in cell c have speeds greater than v_2^{max} .
- 5. Determine the number of trial collisions between a particle in cell c and a neighboring particle by rounding to an integer the expected value

$$N_{tc} = \Gamma_{tc} N_c N_p \Delta t$$

in an unbiased manner, where Δt is the timestep. Here the local trial collision rate is

$$\Gamma_{tc} = \frac{\chi K_c^{max}}{2D},$$

where K_c^{max} is an upper bound for the pre-collision kernel among all candidate pairs. For Maxwell collisions $K_c^{max} = 3\sqrt{k_BT_0/m}/(2\sqrt{\pi})$, and for traditional collisions $K_c^{max} = 3v_{rel}^{(max)}/4$, where $v_{max}^{(rel)} = (v_1^{max} + v_2^{max})$ is a local upper bound on the relative velocity of a colliding pair.

- 6. Perform N_{tc} trial collisions by randomly selecting N_{tc} pairs of particles $i \in \mathcal{L}_1$ and $j \in \mathcal{L}_2$. For each pair, do the following steps if $i \neq j$:
 - (a) Calculate the distance l_{ij} between the centroid of particles i and j, and go to the next pair if $l_{ij} > D$.
 - (b) Calculate the collision kernel $K_{ij}^c = K_c(\boldsymbol{v}_{ij}, \boldsymbol{r}_{ij})$, and go to the next pair if $K_{ij}^c = 0$.
 - (c) Sample a random uniform variate $0 < r \le 1$ and go to the next pair if $K_{ij}^c \le r K_c^{max}$ (note that this step can be skipped in Maxwell I-DSMC since $K_{ij}^c = K_c^{max}$).

(d) Process a stochastic collision between the two particles by updating the particle velocities by sampling the post-collision kernel $P_c(\mathbf{v}'_{ij}; \mathbf{v}_{ij}, \mathbf{r}_{ij})$. For ideal fluids we perform the usual stochastic DSMC collision by randomly rotating \mathbf{v}_{ij} to obtain \mathbf{v}'_{ij} , independent of \mathbf{r}_{ij} .

II. STOCHASTIC HARD SPHERE DYNAMICS

The traditional DSMC fluid has no internal structure and has an ideal gas equation of state (EOS) $p = PV/Nk_BT = 1$, and are thus very compressible. As for the classical hard-sphere fluid, the pressure of fluids with stochastic collisions consists of two parts, an unchanged kinetic contribution that gives the usual ideal-gas pressure $p_{kin} = 1$, and a collisional contribution proportional to the virial $p_{coll} \sim \langle (\boldsymbol{v}_{ij} \cdot \boldsymbol{r}_{ij})' - (\boldsymbol{v}_{ij} \cdot \boldsymbol{r}_{ij}) \rangle_c$, where the average is over stochastic collisions and primes denote post-collisional values. The virial vanishes for collision kernels where velocity updates and positions are uncorrelated leaving only the ideal-gas kinetic contribution. In order to introduce a non-trivial equation of state it is necessary to either give an additional displacement $\Delta \boldsymbol{r}_{ij}$ to the particles that is parallel to \boldsymbol{v}_{ij} , or to bias the momentum exchange $\Delta \boldsymbol{p}_{ij} = m\Delta \boldsymbol{v}_{ij}$ to be (statistically) aligned to \boldsymbol{r}_{ij} . The former approach has already been investigated in the Consistent Boltzmann Algorithm (CBA) [20, 21]. This algorithm was named "consistent" because both the transport coefficients and the equation of state are consistent with those of a hard-sphere fluid to lowest order in density, unlike traditional DSMC which only matches the transport coefficients. However, CBA is not thermodynamically consistent since it modifies the compressibility without affecting the density fluctuations (i.e., the structure of the fluid is still that of a perfect gas).

Here we explore the option of biasing the stochastic momentum exchange based on the position of the colliding particles. What we are trying to emulate through this bias is an effective repulsion between overlapping particles. This repulsion will be maximized if we make Δp_{ij} parallel to r_{ij} , that is, if we use the hard-sphere collision rule $P_c(\mathbf{v}'_{ij}; \mathbf{v}_{ij}, \mathbf{r}_{ij}) = \delta(\mathbf{v}_{ij} - 2v_n \hat{\mathbf{r}}_{ij})$, where $v_n = -\mathbf{v}_{ij} \cdot \hat{\mathbf{r}}_{ij}$ is the normal component of the relative speed. Explicitly, we collide particles as if they are elastic hard spheres of diameter equal to the distance between them at the time of the collision,

$$\boldsymbol{v}_i' = \boldsymbol{v}_i + v_n \hat{\boldsymbol{r}}_{ij}$$

$$\boldsymbol{v}_{j}' = \boldsymbol{v}_{j} - v_{n}\hat{\boldsymbol{r}}_{ij}.$$

Such collisions produce a positive virial only if the particles are approaching each other, i.e., if $v_n > 0$, therefore, we reject collisions among particles that are moving apart, $K_c(\boldsymbol{v}_{ij}, \boldsymbol{r}_{ij}) \sim \Theta(v_n)$,

where Θ denotes the Heaviside function.

To avoid rejection of candidate collision pairs and thus make the algorithm most efficient, it would be ideal if the pre-collision kernel K_c is independent of the relative velocity as for Maxwell collisions. However, without rejection based on the normal v_n or relative v_r speeds, fluctuations of the local temperature T_c would not be consistently coupled to the local pressure. Namely, without rejection the local collisional frequency Γ_{sc} would be independent of T_c and thus the collisional contribution to the pressure $p_c \sim \langle \Delta v_{ij} \cdot r_{ij} \rangle_c \sim \Gamma_{sc} \sqrt{T_c}$ would be $p_c \sim \sqrt{T_c}$ instead of $p_c \sim T_c$, as is required for a fluid with no internal energy. Instead, as for hard spheres, we require that $\Gamma_{sc} \sim \sqrt{T_c}$, which is satisfied if the collision kernel is linear in the magnitude of the relative velocity. For DSMC the collisional rules can be manipulated arbitrarily to obtain the desired transport coefficients, however, for non-ideal fluids thermodynamic requirements eliminate some of the freedom. This important observation has not been taken into account in other algorithms that randomize hard-sphere molecular dynamics [29].

There are two obvious choices for a pre-collision kernel that are linear in the magnitude of the relative velocity. One is to use the relative speed, $K_c \sim v_r$, as in the traditional DSMC algorithm, and the other is to use the hard-sphere pre-collision kernel, $K_c \sim v_n$. We have chosen to make the collision probability linear in the normal speed v_n , specifically, we take $K_c = 3v_n\Theta(v_n)$ to define the Stochastic Hard-Sphere Dynamics (SHSD) fluid. This fluid is simulated by a specific form of the I-DSMC method, in the limit of sufficiently small time steps. Specifically, in the SHSD algorithm we use the following collision kernels in Algorithm 1:

$$K_c = 3v_n\Theta(v_n)$$
 and $K_c^{(max)} = 3v_{max}^{(rel)}$
 $P_c = \delta(\boldsymbol{v}_{ij} - 2v_n\hat{\boldsymbol{r}}_{ij})$
where $v_n = -\boldsymbol{v}_{ij} \cdot \hat{\boldsymbol{r}}_{ij}$.

These choices make the SHSD fluid identical to the one proposed in Ref. [28] for the purposes of proving convergence of a microscopic model to the Navier-Stokes equations. Note that considering particles in neighboring cells as collision partners is essential in SHSD in order to ensure isotropy of the collisional (non-ideal) component of the pressure tensor.

A. Enskog Kinetic Theory

In this section we develop some kinetic equations for the SHSD fluid that are inspired by the Enskog theory of hard-sphere fluids. Remarkeably, it turns out that these sorts of kinetic equations have already been studied in the literature for purely theoretical purposes.

1. BBGKY Hierarchy

The full Bogoliubov-Born-Green-Kirkwood-Yvon (BBGKY) hierarchy of Master equations describing the SHSD fluid is derived in Ref. [28]. Specifically, the evolution of the s-particle distribution function $f_s(t; \mathbf{r}_1, \mathbf{v}_1, \dots, \mathbf{r}_s, \mathbf{v}_s)$ is governed by

$$\frac{\partial f_s}{\partial t} + \sum_{i=1}^s \mathbf{v}_i \cdot \nabla_{\mathbf{r}_i} f_s = \frac{3\chi}{D} \int_0^D dx \int_{\mathcal{R}^3} d\mathbf{v}_j \int_{S_+^2} d\mathbf{e} \quad x^2 \sum_{i=1}^s v_n$$

$$[f_{s+1}(t; \boldsymbol{r}_1, \boldsymbol{v}_1, \dots, \boldsymbol{r}_i, \boldsymbol{v}_i', \dots, \boldsymbol{r}_s, \boldsymbol{v}_s, \boldsymbol{r}_i + x\hat{\boldsymbol{r}}_{ij}, \boldsymbol{v}_j') - f_{s+1}(t; \boldsymbol{r}_1, \boldsymbol{v}_1, \dots, \boldsymbol{r}_i, \boldsymbol{v}_i, \dots, \boldsymbol{r}_s, \boldsymbol{v}_s, \boldsymbol{r}_i - x\hat{\boldsymbol{r}}_{ij}, \boldsymbol{v}_j)],$$
(1)

which takes into account the contribution from collisions of one of the s particles, particle i, with another particle j that is at a distance $r_{ij} = xD$ away, $0 \le x \le 1$. Here S_+^2 denotes the fraction of the unit sphere for which $v_n = -\hat{\boldsymbol{r}}_{ij} \cdot (\boldsymbol{v}_i - \boldsymbol{v}_j) \ge 0$, and $\boldsymbol{v}_i' = \boldsymbol{v}_i + v_n \hat{\boldsymbol{r}}_{ij}$ and $\boldsymbol{v}_j' = \boldsymbol{v}_j - v_n \hat{\boldsymbol{r}}_{ij}$. Just like the BBGKY hierarchy for Hamiltonian fluids, Eqs. (1) are exact, however, they form an infinite unclosed system in which the (s+1)-particle distribution function appears in the equation for the s-particle distribution function. As usual, we need to make an anzatz to truncate and close the system, as we do next.

2. Transport Properties

The hydrodynamics of the SHSD fluid is well-described by a kinetic equation for the singleparticle probability distribution $f(t, \mathbf{r}, \mathbf{v}) \equiv f_1(t; \mathbf{r}, \mathbf{v})$ obtained by making the common molecular chaos assumption about the two-particle distribution function,

$$f_2(t; \mathbf{r}_1, \mathbf{v}_1, \mathbf{r}_2, \mathbf{v}_2) = g_2(\mathbf{r}_1, \mathbf{r}_2; n) f(t, \mathbf{r}_1, \mathbf{v}_1) f(t, \mathbf{r}_2, \mathbf{v}_2),$$

where $g_2(\mathbf{r}_i, \mathbf{r}_j; n)$ is the non-equilibrium pair distribution function that is a functional of the local number density $n(\mathbf{r})$. At global equilibrium $n(\mathbf{r}) = \text{const}$ and $g_2 \equiv g_2(r_{ij})$ depends only on the radial distance once the equilibrium density n and collision rate χ are specified. Substituting the above assumption for f_2 in the first equation of the BBGKY hierarchy (1), we get a stochastic revised Enskog equation of the form studied in Ref. [26],

$$\frac{\partial f(t, \boldsymbol{r}, \boldsymbol{v})}{\partial t} + \boldsymbol{v} \cdot \boldsymbol{\nabla}_{\boldsymbol{r}} f(t, \boldsymbol{r}, \boldsymbol{v}) = \frac{3\chi}{D} \int_{0}^{D} dx \int_{\mathcal{R}^{3}} d\boldsymbol{w} \int_{S_{+}^{2}} d\boldsymbol{e}$$

$$x^{2}\left[g_{2}(\boldsymbol{r},\boldsymbol{r}+x\boldsymbol{e};n)f(t,\boldsymbol{r},\boldsymbol{v}')f(t,\boldsymbol{r}+x\boldsymbol{e},\boldsymbol{w}')-g_{2}(\boldsymbol{r},\boldsymbol{r}-x\boldsymbol{e};n)f(t,\boldsymbol{r},\boldsymbol{v})f(t,\boldsymbol{r}-x\boldsymbol{e},\boldsymbol{w})\right]v_{n},\quad(2)$$

where
$$v_n = -\boldsymbol{e} \cdot (\boldsymbol{v} - \boldsymbol{w}) \ge 0$$
, $\boldsymbol{v}' = \boldsymbol{v} + \boldsymbol{e}v_r$ and $\boldsymbol{w}' = \boldsymbol{w} - \boldsymbol{e}v_r$.

The standard second-order Chapman-Enskog expansion has been carried out for the "stochastic Enskog" equation of the same form as Eq. (2) in Ref. [26], giving estimates of the equation of state (EOS) $p = PV/Nk_BT$, the diffusion coefficient ζ , the shear η and bulk η_B viscosities, and thermal conductivity κ of the SHSD fluid. The expressions in Ref. [26] ultimately express the transport coefficients in terms of various dimensionless integer moments of the pair correlation function $g_2(x = r/D)$, $x_k = \int_0^1 x^k g_2(x) dx$, specifically,

$$p - 1 = 12\phi \chi x_3,\tag{3}$$

$$\zeta/\zeta_0 = \frac{\sqrt{\pi}}{48\phi\chi x_2},\tag{4}$$

$$\eta_B/\eta_0 = \frac{48\phi^2 \chi x_4}{\pi^{3/2}},\tag{5}$$

$$\eta/\eta_0 = \frac{5}{48\sqrt{\pi}\chi x_2} \left(1 + \frac{24\phi\chi x_3}{5}\right)^2 + \frac{3}{5}\eta_B, \text{ and}$$
(6)

$$\kappa/\kappa_0 = \frac{25}{64\sqrt{\pi}\chi x_2} \left(1 + \frac{36\phi\chi x_3}{5}\right)^2 + \frac{3}{2}\eta_B,\tag{7}$$

where $\zeta_0 = D\sqrt{kT/m}$, $\eta_0 = D^{-2}\sqrt{mkT}$ and $\kappa_0 = kD^{-2}\sqrt{kT/m}$ are natural units. These equations are very similar to the ones in the Enskog theory of the hard-sphere fluid except that various coefficients are replaced with moments of $g_2(x)$. In order to use these equations, however, we need to have a good approximation to the pair correlation function, i.e., to the structure of the SHSD fluid. It is important to point out that Eq. (3) is exact as it can be derived directly from the definition of the collisional contribution to the pressure.

B. Pair Correlation Function

In this section we study the structure of the SHSD fluid, theoretically at low densities, and then numerically at higher densities. We find, surpisingly, that there is a thermodynamic correspondence between the stochastic SHSD fluid and a deterministic penetrable-sphere fluid.

1. Low Densities

In order to understand properties of the SHSD fluid as a function of the density ϕ and the cross-section factor χ , we first consider the equilibrium pair correlation function $g_2(r)$ at low densities, where correlations higher than pairwise can be ignored. We consider the cloud of point walkers ij representing the N(N-1)/2 pairs of particles, each at position $\mathbf{r} = \mathbf{r}_i - \mathbf{r}_j$ and with velocity $\mathbf{v} = \mathbf{v}_i - \mathbf{v}_j$. If one of these walkers is closer than D to the origin, $r \leq D$, and is approaching the origin, $v_n > 0$, it reverses its radial speed as a stochastic process with a time-dependent rate $\Gamma = |v_n|\Gamma_0$, where $\Gamma_0 = 3\chi/D$ is the collision frequency. A given walker corresponding to pair ij also undergoes stochastic spatially-unbiased velocity changes with some rate due to the collisions of i with other particles. At low densities we can assume these additional collisions to merely thermalize the velocities to a Maxwell-Boltzmann distribution but not otherwise couple with the radial dependence of the one-particle distribution function $f(\mathbf{v}, \mathbf{r})$ for the N(N-1)/2 walkers. Inside the core $r \leq D$ this distribution of pair walkers satisfies a kinetic equation

$$\frac{\partial f}{\partial t} - v_n \frac{\partial f}{\partial r} = \begin{cases} -\Gamma f & \text{if } v_n \ge 0 \\ \Gamma f & \text{if } v_n < 0 \end{cases} = -\Gamma_0 v_n f.$$

At equilibrium, $\partial f/\partial t = 0$ and v_n cancels on both sides, consistent with choosing collision probability linear in $|v_n|$, giving $\partial f/\partial r = 3\chi D^{-1}f$. At equilibrium, the distribution of the point walkers in phase space ought to be of the separable form $f(\boldsymbol{v}, \boldsymbol{r}) = f(v_n, r) \sim g_2(r) \exp(-mv_n^2/4kT)$, giving $dg_2(r)/dr = 3\chi D^{-1}g_2(r)$ for $r \leq D$ and zero otherwise, with solution

$$g_2(x = r/D) = \begin{cases} \exp[3\chi(x-1)] & \text{for } x \le 1\\ 1 & \text{for } x > 1 \end{cases}$$
 (8)

Indeed, numerical experiments confirmed that at sufficiently low densities the equilibrium g_2 for the SHSD fluid has the exponential form (8) inside the collision core. From statistical mechanics we know that for a deterministic Hamiltonian particle system with a pairwise potential U(r) at low density $g_2^U = \exp[-U(r)/kT]$. Therefore, the low density result (8) is consistent with an effective linear core pair potential

$$U_{eff}(r)/kT = 3\chi(1-x)\Theta(1-x). \tag{9}$$

Note that this repulsive potential is similar to the quadratic core potential used in DPD and strictly vanishes outside of the overlap region, as expected. Also note that the collision parameter χ plays the role that U(0)/kT plays in the system of penetrable spheres interacting with a linear core pairwise potential.

As pointed out earlier, Eq. (3) is exact. At the same time, it is equivalent to the virial theorem for the linear core potential (constant derivative). Therefore, if the pair correlation functions of SHSD and the linear core fluid are truly identical, the pressure of the SHSD fluid is identically equal to that of the corresponding penetrable sphere system. As a consequence, thermodynamic consistency between the structure $(g_2(x))$ and S(k) and equation-of-state for SHSD $(p(\phi))$ is guaranteed to be exact.

2. Equivalence to the Linear Core Penetrable Sphere System

Remarkably, we find numerically that the effective potential (9) can predict exactly $g_2(x)$ at all densities. In fact, we have numerically observed that the SHSD fluid behaves thermodynamically identically to a system of penetrable spheres interacting with a linear core pairwise potential for all ϕ and χ . Figure 1 shows a comparison between the pair correlation function of the SHSD fluid on one hand, and a Monte Carlo calculation using the linear core pair potential on the other, at several densities. Also shown is a numerical solution to the hypernetted chain (HNC) integral equations for the linear core system, inspired by its success for the Gaussian core model [25]. The excellent agreement at all densities permits the use of the HNC result in practical applications, notably the calculation of the transport coefficients via the Enskog-like kinetic theory presented in Section II A 2. We also show the static structure factor S(k) in Fig. 1, and find very good agreement between numerical results and the HNC theory, as expected since S(k) can be expressed as the Fourier transform of $h(r) = g_2(r) - 1$.

Interestingly, in the limit $\chi \to \infty$ the SHSD algorithm reduces to hard-sphere (HS) molecular dynamics. In fact, if the density ϕ is smaller than the freezing point for the HS system, the structure of the SHSD fluid approaches, as χ increases, that of the HS fluid. For higher densities, if χ is sufficiently high, crystallization is observed in SHSD, either to the usual hard-sphere crystals if ϕ is lower than the close-packing density, or if not, to an unusual partially ordered state with multiple occupancy per site, typical of weakly repulsive potentials [30]. Monte Carlo simulations of the linear core penetrable sphere system show identical freezing behavior with SHSD, confirming the surprising equivalence even at non-fluid densities. This points to a conjecture that the (unique) stationary solution to the BBGKY hierarchy (1) is the equilibrium Gibbs distribution,

$$f_s^E = rac{\prod_{i=1}^s M(\boldsymbol{v}_i)}{Z_N} \int_{\boldsymbol{r}_{s+1}} \dots \int_{\boldsymbol{r}_N} \exp \left[-\beta \sum_{i < j} U_{eff}(r_{ij}) \right] d\boldsymbol{r}_{s+1} \dots d\boldsymbol{r}_N,$$

where M is a Maxwellian.

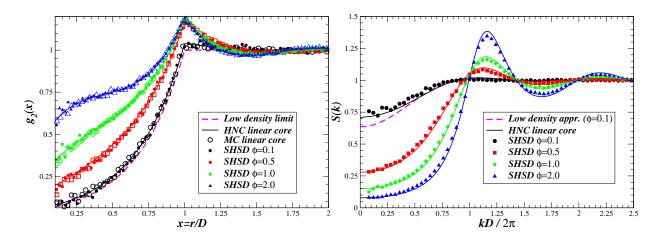


Figure 1: (Left) Equilibrium pair correlation function of the SHSD fluid (solid symbols), compared to MC (open symbols) and numerical solution of the HNC equations (solid lines) for the linear core system, at various densities and $\chi = 1$. The low-density approximation corresponding to Eq. (8) is also shown. (Right) The corresponding static structure factors from SHSD simulations (solid symbols) and HNC calculations (solid lines).

III. RESULTS

In this Section we perform several numerical experiments with the SHSD algorithm. Firstly, we compare the theoretical predictions for the transport properties of the SHSD fluid based on the HNC theory for the linear core penetrable sphere system with results from particle simulations. We then compute dynamic structure factors and compare them to predictions of fluctuating hydrodynamics. Finally, we study the motion of a Brownian bead suspended in an SHSD fluid.

A. Transport Coefficients

The equation of state of the SHSD fluid for a given χ is $P = p(\phi)Nk_BT/V$, where $p(\phi)$ is given in Eq. (3). According to statistical mechanics, the structure factor at the origin, is equal to the isothermal compressibility, that is,

$$S_0 = S(\omega = 0, k = 0) = \tilde{c}_T^{-2} = (p + \phi dp/d\phi)^{-1}$$

where $c_T = \tilde{c}_T \sqrt{k_B T/m}$ is the isothermal speed of sound. In the inset in the top part of Fig. 2, we directly demonstrate the thermodynamic consistency of SHSD by comparing the compressibility calculated through numerical differentiation of the pressure, to the structure factor at the origin. The pressure is easily measured in the particle simulations by keeping track of the total collisional momentum exchange during a long period. The structure factor is obtained through a temporal

average of a Fast Fourier Transform approximation to the discrete Fourier Transform of the particle positions $\|\sum_i \exp(-i\mathbf{k}\cdot\mathbf{r}_i)\|^2$. The value S(k=0) is estimated by fitting a parabolic dependence for small k and extrapolating to k=0.

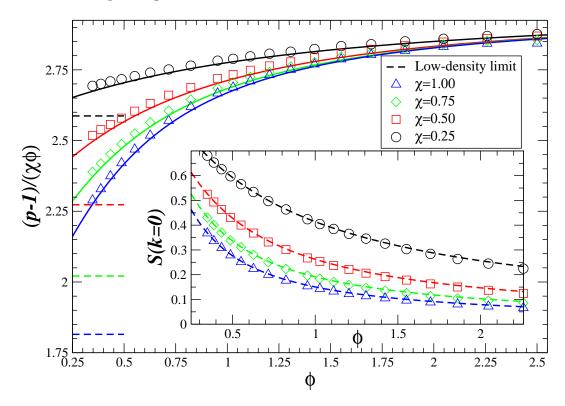


Figure 2: Normalized equation of state $(p-1)/(\chi\phi)$ for the SHSD fluid at several cross-section factors χ (different symbols) compared to theoretical predictions based on the virial theorem (3) with the HNC approximation to $g_2(x)$ (solid lines). The inset compares the compressibility (pressure derivative, dashed lines) to the structure factor at the origin $S(k \to 0)$ (symbols), measured using a direct Fourier transform of the particle positions for small k and extrapolating to k = 0.

Having established that the HNC closure provides an excellent approximation $g_2^{(HNC)}(x)$ for the pair correlation function of the SHSD fluid, we can obtain estimates for the transport coefficients by calculating the first four moments of $g_2^{(HNC)}(x)$ and substituting them in the results of the Enskog kinetic theory presented in Section II A 2. In Figure 3 we compare the theoretical predictions for the diffusion coefficient ζ and the viscosity η to the ones directly calculated from SHSD particle simulations. We measure ζ directly from the average mean square displacement of the particles. We estimate η by calculating the mean flow rate in Poiseuille parabolic flow between two thermal hard walls due to an applied constant force on the particles². Surprisingly, good agreement is found for

² Similar results are obtained by calculating the viscous contributions to the kinetic and collisional stress tensor

the shear viscosity at all densities. Similar matching was observed for the thermal conductivity κ . The corresponding results for the diffusion coefficient show significant ($\sim 25\%$) deviations for the self-diffusion coefficient at higher densities because of corrections due to higher-order correlations.

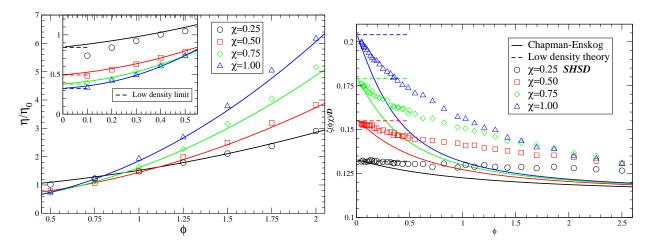


Figure 3: Comparison between numerical results for SHSD at several collision frequencies (different symbols) with predictions based on the stochastic Enskog equation using the HNC approximation to $g_2(x)$ (solid lines). The low-density approximations are also indicated (dashed lines). (Left) The normalized shear viscosity η/η_0 at high and low densities (inset), as measured using an externally-forced Poiseuille flow. There are significant corrections (Knudsen regime) for large mean free paths (i.e., at low densities and low collision rates). (Right) The normalized diffusion coefficient $\zeta/(\chi\phi\zeta_0)$, as measured from the mean square displacement of the particles.

B. Dynamic Structure Factors

The hydrodynamics of the spontaneous thermal fluctuations in the SHSD fluid is expected to be described by the Landau-Lifshitz Navier-Stokes (LLNS) equations linearized around a reference equilibrium state $U_0 = (\rho_0, v_0 = \mathbf{0}, T_0)$ [31, 32]. For the SHSD fluid the linearized equation of state is

$$P = p(\phi) \frac{Nk_B T}{V} \approx (p_0 + \tilde{c}_T^2 \frac{\rho}{\rho_0} + p_0 \frac{T}{T_0}) \rho_0 c_0^2,$$

and there is no internal energy contribution to the energy density,

$$e \approx \frac{3}{2} \frac{Nk_B T}{V} = e_0 + \rho c_v T_0 + \rho_0 c_v T,$$

in non-equilibrium simulations of Couette shear flow. This kind of calculation additionally gives the split in the viscosity between kinetic and collisional contributions.

where $p_0 = p(\phi_0)$, $c_0 = k_B T/m$, and $c_v = 3k_B/2m$, giving an adiabatic speed of sound $c_s = \tilde{c}_s c_0$, where $\tilde{c}_s^2 = \tilde{c}_T^2 + 2p^2/3$. Therefore, for one-dimensional flows the LLNS equations for the fluctuating field $U = (\rho_0 + \rho, v, T_0 + T)$ take the form

$$\begin{bmatrix} \partial_{t}\rho \\ \partial_{t}v \\ \partial_{t}T \end{bmatrix} = -\frac{\partial}{\partial x} \begin{bmatrix} \rho_{0}v \\ c_{T}^{2}\rho_{0}^{-1}\rho + p_{0}c_{0}^{2}T_{0}^{-1}T \\ p_{0}c_{0}^{2}c_{v}^{-1}v \end{bmatrix} + \frac{\partial}{\partial x} \begin{bmatrix} 0 \\ \rho_{0}^{-1}\eta_{0}v_{x} \\ \rho_{0}^{-1}c_{v}^{-1}\kappa_{0}T_{x} \end{bmatrix} + \frac{\partial}{\partial x} \begin{bmatrix} 0 \\ \rho_{0}^{-1}\sqrt{2\eta_{0}k_{B}T_{0}}W^{(v)} \\ \rho_{0}^{-1}c_{v}^{-1}T_{0}\sqrt{2\kappa_{0}k_{B}}W^{(T)} \end{bmatrix},$$
(10)

where $W^{(v)}$ and $W^{(T)}$ are independent spatio-temporal white noise Gaussian fields.

By solving these equations in the Fourier wavevector-frequency domain for $\widehat{U}(k,\omega)$ and performing an ensemble average over the fluctuating stresses we can obtain the spatio-temporal correlations (covariance) of the fluctuating fields. We express these correlations in terms of the 3×3 symmetric positive-definite hydrodynamic structure factor matrix $S_H(k,\omega) = \left\langle \widehat{U}\widehat{U}^* \right\rangle$ [5]. The hydrostatic structure factor matrix $S_H(k)$ is obtained by integrating $S_H(k,\omega)$ over all frequencies,

$$\mathbf{S}_{H}(k) = \begin{bmatrix} \rho_{0}c_{T}^{-2}k_{B}T_{0} & 0 & 0\\ 0 & \rho_{0}^{-1}k_{B}T_{0} & 0\\ 0 & 0 & \rho_{0}^{-1}c_{v}^{-1}k_{B}T_{0}^{2} \end{bmatrix}.$$

$$(11)$$

We use $\mathbf{S}_H(k)$ for an ideal gas (i.e., for $p_0 = 1$, $\tilde{c}_T = 1$) to non-dimensionalize $\mathbf{S}_H(k,\omega)$, for example, we express the spatio-temporal cross-correlation between density and velocity through the dimensionless hydrodynamic structure factor

$$S_{\rho,v}(k,\omega) = \left(\rho_0 c_0^{-2} k_B T_0\right)^{-\frac{1}{2}} \left(\rho_0^{-1} k_B T_0\right)^{-\frac{1}{2}} \left\langle \hat{\rho}(k,\omega) \hat{v}^*(k,\omega) \right\rangle.$$

For the non-ideal SHSD fluid the density fluctuations have a spectrum

$$S_{\rho}(k) = (\rho_0 c_0^{-2} k_B T_0)^{-1} \langle \hat{\rho}(k) \hat{\rho}^{\star}(k) \rangle = \tilde{c}_T^{-2},$$

which only captures the small k behavior of the full (particle) structure factor S(k) (see Fig. 1), as expected of a continuum theory that does not account for the structure of the fluid. Typically only the density-density dynamic structure factor is considered because it is accessible experimentally via light scattering measurements and thus most familiar. However, in order to fully access the validity of the full LLNS system one should examine the dynamic correlations among all pairs of variables. The off-diagonal elements of the static structure factor matrix $S_H(k)$ vanish because the primitive hydrodynamic variables are instantaneously uncorrelated, however, they have non-trivial

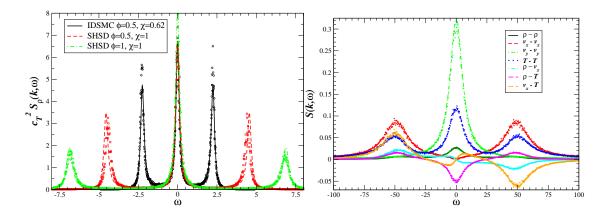


Figure 4: Hydrodynamic structure factors as predicted by the LLNS equations (solid lines) and computed from particle simulations (symbols, collision diameter D=0.04). For the SHSD fluid we obtain the transport coefficients from the Enskog theory with the HNC approximation to g_2 , while for the Maxwell I-DSMC fluid we numerically estimated the viscosity and thermal conductivity. (Left) Normalized density fluctuations $\tilde{c}_T^2 S_{\rho}(k,\omega)$ for $k\approx 1.75$ for an ideal Maxwell I-DSMC ($\phi=0.5,\,\chi=0.62$) and two non-ideal SHSD ($\phi=0.5$ and $\phi=1,\,\chi=1$) fluids, showing the shifting of the side Brillouin peaks to higher frequencies as the compressibility is reduced. (Right) Selected diagonal and off-diagonal elements of the non-dimensionalized hydrodynamic structure factor matrix for a larger $k\approx 12.6$ for an SHSD fluid at $\phi=1,\,\chi=1$, giving a complete picture of the spatio-temporal correlations between all of the different pairs of hydrodynamic variables.

dynamic correlations visible in the off-diagonal elements of the dynamic structure factor matrix $S_H(k,\omega)$.

In Fig. 4 we compare theoretical and numerical results the hydrodynamic structure factors for the SHSD fluid with $\chi=1$ at two densities for a small and a medium k value $[kD/(2\pi)\approx 0.01 \text{ and } 0.07]$. In this figure we show selected elements of $S_H(k,\omega)$ as predicted by the analytical solution to Eqs. (10) with parameters obtained by using the HNC approximation to g_2 in the Enskog kinetic theory presented in Section II A 2. Therefore, for SHSD the theoretical calculations of $S_H(k,\omega)$ do not use any numerical inputs from the particle runs. We also show hydrodynamic structure factors obtained from particle simulations in a quasi-one-dimensional setup in which the simulation cell was periodic and long along the x axis, and divided into hydrodynamic cells of length 5D. Finite-volume averages of the hydrodynamic conserved variables were then calculated for each cell every 10 time steps and a Fast Fourier Transform used to obtain hydrodynamic structure factors for several wavenumbers. Figure 4 shows very good agreement between theory and numerics, and clearly shows the shifting of the two symmetric Brillouin peaks at $\omega \approx c_s k$ toward higher frequencies as the compressibility of the SHSD fluid is reduced and the speed of

sound increased. The width of the central Rayleigh peak at $\omega = 0$ is also well-predicted for a wide range of k values, demonstrating that the SHSD fluid shows the expected fluctuating hydrodynamic behavior.

C. Brownian Walker VACF

As an illustration of the correct hydrodynamic behavior of the SHSD fluid and the significance of compressibility, we study the velocity autocorrelation function (VACF) $C(t) = \langle v_x(0)v_x(t)\rangle$ for a single neutrally-buoyant hard sphere of mass M and radius R suspended in an SHSD fluid of mass density ρ . This problem is relevant to the modeling of polymer chains or (nano)colloids in solution, and led to the discovery of a long power-law tail in C(t) [33, 34]. Here the fluid particles interact via stochastic collisions, exactly as in I-DSMC. The interaction between fluid particles and the bead is treated as if the SHSD particles are hard spheres of diameter D_s , chosen to be somewhat smaller than their interaction diameter with other fluid particles (specifically, we use $D_s = D/4$) for computational efficiency reasons, using an event-driven algorithm [35]. Upon collision with the bead the relative velocity of the fluid particle is reversed in order to provide a no-slip condition at the surface of the suspended sphere [33, 35] (slip boundaries give qualitatively identical results). For comparison, an ideal I-DSMC fluid of comparable viscosity is also simulated.

Theoretically, C(t) has been calculated from the linearized (compressible) fluctuating Navier-Stokes (NS) equations [33]. The results are analytically complex even in the Laplace domain, however, at short times an inviscid compressible approximation applies. At large times the compressibility does not play a role and the incompressible NS equations can be used to predict the long-time tail. At short times, $t < t_c = 2R/c_s$, the major effect of compressibility is that sound waves generated by the motion of the suspended particle carry away a fraction of the momentum, so that the VACF quickly decays from its initial value C(0) = kT/M to $C(t_c) \approx kT/M_{eff}$, where $M_{eff} = M + 2\pi R^3 \rho/3$. At long times, $t > t_{visc} = 4\rho R_H^2/3\eta$, the VACF decays as in an incompressible fluid, with an asymptotic power-law tail $(kT/m)(8\sqrt{3\pi})^{-1}(t/t_{visc})^{-3/2}$, in disagreement with predictions based on the Langevin equation (Brownian dynamics), $C(t) = (kT/m) \exp(-6\pi R_H \eta t/m)$. We have estimated the effective (hydrodynamic) colloid radius R_H from numerical measurements of the Stokes friction force $F = -6\pi R_H \eta v$ and found it to be somewhat larger than the hard-core collision radius $R + D_s/2$.

In Fig. 5 numerical results for the VACF for an I-DSMC fluid and an SHSD fluid at two different densities are compared to the theoretical predictions. It is seen, as predicted, that the

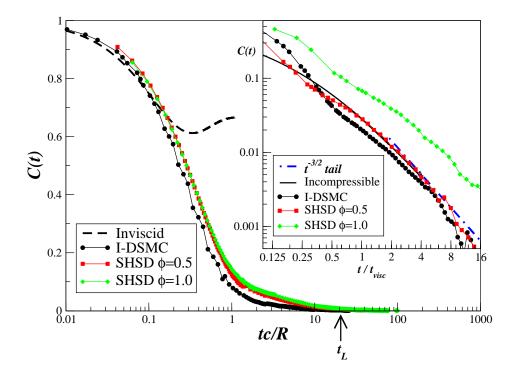


Figure 5: The velocity autocorrelation function for a neutrally buoyant hard sphere suspended in a non-ideal SHSD ($\chi=1$) fluid at two densities (symbols), as well as an ideal Maxwell I-DSMC fluid ($\phi=0.5, \chi=0.62$, symbols), at short and long times (inset). For the more compressible (less viscous) fluids the long time tails are statistically measurable only up to $t/t_{visc}\approx 5$. The theoretical predictions based on the inviscid, for short times, or incompressible, for long times, Navier-Stokes equations are also shown (lines). The diameter of the nano-colloidal particle is only 2.5D, although we have performed simulations using larger spheres as well with very similar (but less accurate) results. Since periodic boundary conditions were used we only show the tail up to about the time at which sound waves generated by its periodic images reach the particle, $t_L=L/c_s$.

compressibility or the sound speed c_s , determines the early decay of the VACF. The exponent of the power-law decay at large times is also in agreement with the hydrodynamic predictions. The coefficient of the VACF tail agrees reasonably well with the hydrodynamic prediction for the less dense fluids, however, there is a significant deviation of the coefficient for the densest fluids, perhaps due to ordering of the fluid around the suspended sphere, not accounted for in continuum theory. In order to study this discrepancy in further detail one would need to perform simulations with a much larger bead. This is prohibitively expensive with the serial event-driven algorithm used here [35] and requires either parallelizing the code or using a hybrid particle-continuum method [36], which we leave for future work.

IV. CONCLUSIONS

We have sucessfully generalized the traditional DSMC algorithm for simulating rare gas flows to flows of dense non-ideal fluids. Constructing such a thermodynamically-consistent Stochastic Hard Sphere Dynamics (SHSD) algorithm required first eliminating the grid artifacts from traditional DSMC. These artifacts are small in traditional DSMC simulations because the collisional cell size is kept significantly smaller than the mean free path [37], but become pronounced when collision-dominated flows are simulated because the collisional-stress tensor is not isotropic. Our Isotropic DSMC (I-DSMC) method is a grid free DSMC variant with pairwise spherically-symmetric stochastic interactions between the particles, just as classical fluids simulated by molecular dynamics (MD) use a pairwise spherically-symmetric deterministic interaction potential. The I-DSMC method can therefore be viewed as a transition from the DSMC method, suitable for rarefied flows, to the MD method, suitable for simulating dense liquids (and solids).

It has long been apparent that manipulating the stochastic collision rules in DSMC can lead to a wide range of fluid models, including non-ideal ones [20, 38]. It has also been realized that DSMC, as a kinetic Monte Carlo method, is not limited to solving the Boltzmann equation but can be generalized to Enskog-like kinetic equations [22, 23]. However, what has been so far elusive is to construct a non-ideal stochastic fluid model that is thermodynamically consistent, meaning that the fluid structure and the equation of state are consistent with each other as required by statistical mechanics. We overcame this hurdle here by constructing stochastic collision kernels in I-DSMC to be as close as possible to those of the classical hard-sphere deterministic system. Thus, in the SHSD algorithm randomly chosen pairs of approaching and overlapping particles undergo collisions as if they were hard spheres of variable diameter.

We demonstrated the consistent thermodynamic behavior of the SHSD system by observing that it has identical structure and thermodynamic properties to a Hamiltonian system of penetrable spheres interacting with a linear core potential, even up to solid densities. We found that at fluid densities the pair correlation function $g_2(r)$ of the linear core system is well-described by the approximate HNC closure, enabling us to obtain moments of $g_2(r)$. These moments were then used as inputs in a modified Chapman-Enskog calculation to obtain excellent estimates of the equation of state and transport coefficients of the SHSD fluid over a wide range of densities. We do not yet have a complete theoretical understanding of our surprising finding that the SHSD system behaves thermodynamically identically to the linear core system. An important open question remains whether by choosing a different collision kernel one can obtain stochastic fluids corresponding to

Hamiltonian systems of penetrable spheres interacting with effective pair potentials $U_{eff}(r)$ other than the linear core potential.

The SHSD algorithm is similar in nature to DPD and has a similar computational complexity. The essential difference is that DPD has a continuous-time formulation (a system of stochastic ODEs), where as the SHSD dynamics is discontinuous in time (Master Equation). This is similar to the difference between MD for continuous potentials and discontinuous potentials. Just as DSMC is a stochastic alternative to hard-sphere MD for low-density gases, SHSD is a stochastic modification of hard-sphere MD for dense gases. On the other hand, DPD is a modification of MD for smooth potentials to allow for larger time-steps and a conservative thermostat.

A limitation of SHSD is that for reasonable values of the collision frequency ($\chi \sim 1$) and density ($\phi \sim 1$) the fluid is still relatively compressible compared to a dense liquid, $S(k=0) = \tilde{c}_T^{-2} > 0.1$. Indicative of this is that the diffusion coefficient is large relative to the viscosity as it is in typical DPD simulations, so that the Schmidt number $S_c = \eta(\rho\zeta)^{-1}$ is less than 10 instead of being on the order of 100-1000. Achieving higher \tilde{c}_T or S_c requires high collision rates (for example, $\chi \sim 10^4$ is used in Ref. [39]) and appropriately smaller time steps to ensure that there is at most one collision per particle per time step, and this requires a similar computational effort as in hard-sphere molecular dynamics at a comparable density. At low and moderate gas densities the SHSD algorithm is not as efficient as DSMC at a comparable collision rate. However, for a wide range of compressibilities, SHSD is several times faster than the alternative deterministic Event-Driven MD (EDMD) for hard spheres [24, 40]. Furthermore, SHSD has several important advantages over EDMD, in addition to its simplicity:

- 1. SHSD has several controllable parameters that can be used to change the transport coefficients and compressibility, notably the usual density ϕ but also the cross-section factor χ and others³, while EDMD only has density.
- 2. SHSD is time-driven rather than event-driven thus allowing for easy parallelization.
- 3. SHSD can be more easily coupled to continuum hydrodynamic solvers, just like ideal-gas DSMC [41] and DPD [42, 43]. Strongly-structured particle systems, such as fluids with strong interparticle repulsion (e.g., hard spheres), are more difficult to couple to hydrodynamic

³ For example, one can combine rejection-free Maxwell collisions with hard-sphere collisions in order to tune the viscosity without affecting the compressibility. The efficiency is significantly enhanced when the fraction of accepted collisions is increased, however, the compressibility is also increased at a comparable collision rate.

solvers [44] than ideal fluids, such as MPCD or (I-)DSMC, or weakly-structured fluids, such as DPD or SHSD fluids.

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Appendix A: I-DSMC AT LOW DENSITIES

The I-DSMC algorithm was designed with dense fluid flows in mind, that is, collision-dominated flows where the mean free path is comparable to the collision range. In the traditional DSMC algorithm the collision range is the cell length L_c , and in I-DSMC the collision range is the particle diameter D.

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